A solution to the density matrix equation

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The authors derive a solution for the density matrix of an electron moving in a spatially periodic field. The solution corresponds to a mixed state of the system.

Consider the motion of a one-dimensional electron in a periodic field described by the Hamiltonian

$$\hat{H} = -\frac{\hbar^2}{2m}\frac{\mathrm{d}^2}{\mathrm{d}x^2} + U\sin kx,$$

where U and k are arbitrary real constants.

The corresponding equation for the density matrix can be written as follows:¹

$$i\hbar \frac{\partial}{\partial t}\rho(x, x', t) = (\hat{H} - \hat{H}'^*)\rho(x, x', t);$$
(1)

where H' is the same operator acting on x'. It is easy to verify by direct substitution that equation (1) has an exact, steadystate solution:

$$\rho(x, x') = C \exp\left\{-\frac{m}{\alpha} \left(\frac{2}{\hbar k}\right)^2 \cos\left[\frac{k}{2} (x - x')\right] - \alpha U \sin\left[\frac{k}{2} (x + x')\right]\right\}; \qquad (2)$$

where C is a normalization constant and α is an arbitrary constant.

Let the one-dimensional motion occur in a segment $x \in [0,L]$. We impose the periodic boundary conditions

$$\rho(x+L, x') = \rho(x, x'), \quad \rho(x, x'+L) = \rho(x, x'). \quad (3)$$

Substituting (2) into (3) one finds that the boundary conditions are satisfied only if

$$k = \frac{4\pi}{L}v, \quad v = 1, 2, 3, \dots$$
 (4)

The normalization constant C is determined by the condition

$$\int_{0}^{L} \rho(x, x) \, \mathrm{d}x = 1$$

whence

$$C = \exp\left[\frac{m}{\alpha}\left(\frac{2}{\hbar k}\right)^{2}\right] \left[\int_{0}^{L} \exp\left(-\alpha U \sin kx\right) dx\right]^{-1}$$
$$= \widetilde{C} \exp\left[\frac{m}{\alpha}\left(\frac{2}{\hbar k}\right)^{2}\right],$$

where

$$\widetilde{C} = \left[\int_{0}^{L} \exp\left(-\alpha U \sin kx\right) \mathrm{d}x\right]^{-1}.$$
(5)

For $\rho(x, x')$ we find, ultimately,

$$\rho(x, x') = \widetilde{C} \exp\left[-\frac{m}{\alpha} \left(\frac{2}{\hbar k}\right)^2 \left\{\cos\left[\frac{k}{2} (x - x')\right] - 1\right\} - \alpha U \sin\left[\frac{k}{2} (x + x')\right]\right].$$
(6)

Let us now derive the Wigner function corresponding to solution (6). We proceed from the general formula²

$$f(p, x) = \int_{0}^{\infty} \rho\left(x + \frac{t}{2}, x - \frac{t}{2}\right) \exp\left(-\frac{ipt}{\hbar}\right) \frac{dt}{2\pi\hbar}.$$
 (7)

Substituting (6) into (7) we obtain

$$f(p, x) = \overline{C} \exp\left(-\alpha U \sin kx\right) \\ \times \exp\left[-\frac{m}{\alpha} \left(\frac{2}{\hbar k}\right)^2\right] \sum_{n=-\infty}^{\infty} I_n\left(\frac{m}{\alpha} \left(\frac{2}{\hbar k}\right)^2\right) \delta\left(p - \frac{n\hbar k}{2}\right),$$
(8)

where $I_n(z)$ is the modified Bessel function of order n; $\delta(z)$ is the delta function; $n = 0, \pm 1, \pm 2,...$. One can verify by direct substitution that the normalization condition is satisfied:²

$$\int f(p, x) \, \mathrm{d}p \, \mathrm{d}x = 1$$

From expression (8) we can solve for the mean value of the electron momentum

$$\langle p \rangle = 0$$

Analogously, we find using (8) that

$$\left< \frac{p^2}{2m} \right> = \frac{1}{2\alpha} \ .$$

The classical limit of f(p, x) can be easily obtained by taking the $\hbar \rightarrow 0$ limit of (8):

$$f(p, x) = \widetilde{C} \left(\frac{\alpha}{2\pi m}\right)^{1/2} \exp\left(-\frac{\alpha p^2}{2m} - \alpha U \sin kx\right).$$
(9)

Note that if parameter α^{-1} is interpreted as a temperature T (in energy units), then solution (9) transforms into the classical Gibbs distribution for a particle in a field $U \sin kx$:

$$f(p, x) = \frac{\widetilde{C}}{(2\pi mT)^{1/2}} \exp\left(-\frac{H(p, x)}{T}\right).$$

Recall that the C of this formula is given by equation (5).

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¹L. D. Landau and E. M. Lifshits, *Quantum Mechanics: Non-Relativistic Theory*, 2nd ed., Pergamon Press, Oxford, 1962 [Russ. original, Nauka, M., 1963].

² R. Feynmann, Statistical Mechanics, Benjamin, N. Y., 1972 [Russ. transl., Mir, M., 1975].